# Advanced Computational Fluid Dynamics AA215A Lecture 4 

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#### Abstract

Lecture 4 covers analysis of the equations of gas dynamics


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Chapter 1
Analysis of the equations of gas dynamics

## Lecture 4

## Analysis of the equations of gas dynamics

### 1.1 Coordinate Transformations

In order to calculate solutions for flows in complex geometric domains, it is often useful to introduce body-fitted coordinates through global, or, as in the case of isoparametric elements, local transformations. With the body now coinciding with a coordinate surface, it is much easier to enforce the boundary conditions accurately. Suppose that the mapping to computational coordinates $\left(\xi_{1}, \xi_{2}, \xi_{3}\right)$ is defined by the transformation matrices

$$
\begin{equation*}
K_{i j}=\frac{\partial x_{i}}{\partial \xi_{j}}, \quad K_{i j}^{-1}=\frac{\partial \xi_{i}}{\partial x_{j}}, \quad J=\operatorname{det}(K) \tag{1.1}
\end{equation*}
$$

The Navier-Stokes equations (?? -?? ) become

$$
\begin{equation*}
\frac{\partial}{\partial t}(J w)+\frac{\partial}{\partial \xi_{i}} F_{i}(w)=0 \tag{1.2}
\end{equation*}
$$

Here the transformed fluxes are

$$
\begin{equation*}
F_{i}=S_{i j} f_{j} \tag{1.3}
\end{equation*}
$$

where

$$
\begin{equation*}
S=J K^{-1} \tag{1.4}
\end{equation*}
$$

The elements of $S$ are the cofactors of $K$, and in a finite volume discretization they are just the face areas of the computational cells projected in the $x_{1}, x_{2}$, and $x_{3}$ directions. Using the permutation tensor $\epsilon_{i j k}$ we can express the elements of $S$ as

$$
\begin{equation*}
S_{i j}=\frac{1}{2} \epsilon_{j p q} \epsilon_{i r s} \frac{\partial x_{p}}{\partial \xi_{r}} \frac{\partial x_{q}}{\partial \xi_{s}} \tag{1.5}
\end{equation*}
$$

Then

$$
\begin{align*}
\frac{\partial}{\partial \xi_{i}} S_{i j} & =\frac{1}{2} \epsilon_{j p q} \epsilon_{i r s}\left(\frac{\partial^{2} x_{p}}{\partial \xi_{r} \partial \xi_{i}} \frac{\partial x_{q}}{\partial \xi_{s}}+\frac{\partial x_{p}}{\partial \xi_{r}} \frac{\partial^{2} x_{q}}{\partial \xi_{s} \partial \xi_{i}}\right) \\
& =0 \tag{1.6}
\end{align*}
$$

Defining scaled contravariant velocity components as

$$
\begin{equation*}
U_{i}=S_{i j} u_{j} \tag{1.7}
\end{equation*}
$$

the flux formulas may be expanded as

$$
F_{i}=\left\{\begin{array}{c}
\rho U_{i}  \tag{1.8}\\
\rho U_{i} u_{1}+S_{i 1} p \\
\rho U_{i} u_{2}+S_{i 2} p \\
\rho U_{i} u_{3}+S_{i 3} p \\
\rho U_{i} H
\end{array}\right\} .
$$

If we choose a coordinate system so that the boundary is at $\xi_{l}=0$, the wall boundary condition for inviscid flow is now

$$
\begin{equation*}
U_{l}=0 . \tag{1.9}
\end{equation*}
$$

### 1.2 Analysis of the equations of gas dynamics: the Jacobian matrices

The Euler equations for the three-dimensional flow of an inviscid gas can be written in integral form, using the summation convention, as

$$
\frac{d}{d t} \int_{\Sigma} \boldsymbol{w} d V+\int_{d \Sigma} \boldsymbol{f}_{i} n_{i} d S=0
$$

where $\Sigma$ is the domain, $d \Sigma$ its boundary, $\vec{n}$ the normal to the boundary, and $d V$ and $d S$ are the volume and area elements. Let $x_{i}, u_{i}, \rho, p, E$ and $H$ denote the Cartesian coordinates velocity, density, pressure, energy and enthalpy. In differential form

$$
\begin{equation*}
\frac{\partial \boldsymbol{w}}{\partial t}+\frac{\partial}{\partial x_{i}} \boldsymbol{f}_{i}(\boldsymbol{w})=0 \tag{1.10}
\end{equation*}
$$

where the state and flux vectors are

$$
\boldsymbol{w}=\rho\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
u_{3} \\
E
\end{array}\right], \quad \boldsymbol{f}_{i}=\rho u_{i}\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
u_{3} \\
H
\end{array}\right]+p\left[\begin{array}{c}
0 \\
\delta_{i 1} \\
\delta_{i 2} \\
\delta_{i 3} \\
0
\end{array}\right]
$$

Also,

$$
p=(\gamma-1) \rho\left(E-\frac{u^{2}}{2}\right), \quad H=E+\frac{p}{\rho}=\frac{c^{2}}{\gamma-1}+\frac{u^{2}}{2}
$$

where $u$ is the speed and $c$ is the speed of sound

$$
u^{2}=u_{i}^{2}, c^{2}=\frac{\gamma p}{\rho}
$$

Let $m_{i}$ and $e$ denote the momentum components and total energy,

$$
m_{i}=\rho u_{i}, e=\rho E=\frac{p}{\gamma-1}+\frac{m_{i}^{2}}{2 \rho}
$$

Then $\boldsymbol{w}$ and $f$ can be expressed as

$$
\boldsymbol{w}=\left[\begin{array}{c}
\rho  \tag{1.11}\\
m_{1} \\
m_{2} \\
m_{3} \\
e
\end{array}\right], \quad \boldsymbol{f}_{i}=u_{i}\left[\begin{array}{c}
\rho \\
m_{1} \\
m_{2} \\
m_{3} \\
e
\end{array}\right]+p\left[\begin{array}{c}
0 \\
\delta_{i 1} \\
\delta_{i 2} \\
\delta_{i 3} \\
u_{i}
\end{array}\right]
$$

In a finite volume scheme the flux needs to be calculated across the interface between each pair of cells. Denoting the face normal and area by $n_{i}$ and $S$, the flux is $F S$ where

$$
\boldsymbol{f}=n_{i} \boldsymbol{f}_{i}
$$

This can be expressed in terms of the conservative variables $\boldsymbol{w}$ as

$$
\boldsymbol{f}=u_{n}\left[\begin{array}{c}
\rho  \tag{1.12}\\
m_{1} \\
m_{2} \\
m_{3} \\
e
\end{array}\right]+p\left[\begin{array}{c}
0 \\
n_{1} \\
n_{2} \\
n_{3} \\
u_{n}
\end{array}\right]
$$

where $u_{n}$ is the normal velocity

$$
u_{n}=n_{i} u_{i}=\frac{n_{i} m_{i}}{\rho}
$$

Also

$$
p=(\gamma-1)\left(e-\frac{m_{i}^{2}}{\rho}\right)
$$

In smooth regions of the flow the equations can also be written in quasilinear form as

$$
\frac{\partial \boldsymbol{w}}{\partial t}+A_{i} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0
$$

where $A_{i}$ are the Jacobian matrices

$$
A_{i}=\frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}},
$$

The composite Jacobian matrix at a face with normal vector $\vec{n}$ is

$$
A=A_{i} n_{i}=\frac{\partial \boldsymbol{f}}{\partial \boldsymbol{w}}
$$

All the entries in $\boldsymbol{f}_{i}$ and $\boldsymbol{f}$ are homogenous of degree 1 in the conservative variables $\boldsymbol{w}$. It follows that $\boldsymbol{f}_{i}$ and $\boldsymbol{f}$ satisfy the identities

$$
\boldsymbol{f}_{i}=A_{i} \boldsymbol{w}, \boldsymbol{f}=A \boldsymbol{w}
$$

This is the consequence of the fact that if a quantity $q$ can be expressed in terms of the components of a vector $\boldsymbol{w}$ as

$$
q=\prod_{j} \boldsymbol{w}_{j}^{\alpha_{j}}, \text { where } \sum_{j} \alpha_{j}=\alpha,
$$

then

$$
\begin{align*}
\sum_{i} \frac{\partial q}{\partial \boldsymbol{w}_{i}} \boldsymbol{w}_{i} & =\sum_{i} \alpha_{i} \prod_{j} \boldsymbol{w}_{j}^{\alpha_{j}}  \tag{1.13}\\
& =\sum_{i} \alpha_{i} q \\
& =\alpha q
\end{align*}
$$

In order to evaluate $A$ note that $\frac{\partial u_{n}}{\partial w}$ and $\frac{\partial p}{\partial w}$ are row vectors

$$
\frac{\partial u_{n}}{\partial \boldsymbol{w}}=\frac{1}{\rho}\left[-u_{n}, n_{1}, n_{2}, n_{3}, 0\right]
$$

and

$$
\frac{\partial p}{\partial \boldsymbol{w}}=(\gamma-1)\left[\frac{u^{2}}{2},-u_{1},-u_{2},-u_{3}, 1\right]
$$

Accordingly

$$
\frac{\partial f}{\partial \boldsymbol{w}}=\frac{\partial}{\partial \boldsymbol{w}}\left(u_{n} \boldsymbol{w}\right)+\left[\begin{array}{c}
0  \tag{1.14}\\
n_{1} \\
n_{2} \\
n_{3} \\
u_{n}
\end{array}\right]\left[\frac{\partial p}{\partial \boldsymbol{w}}\right]+\left[\begin{array}{l}
0 \\
0 \\
0 \\
0 \\
p
\end{array}\right]\left[\frac{\partial u_{n}}{\partial \boldsymbol{w}}\right]
$$

Then the Jacobian matrix can be assembled as the sum of a diagonal matrix and two outer products of rank 1 ,

$$
\begin{aligned}
A & =u_{n} I+\rho\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
u_{3} \\
H
\end{array}\right]\left[\frac{\partial u_{n}}{\partial w}\right]+\left[\begin{array}{c}
0 \\
n_{1} \\
n_{2} \\
n_{3} \\
u_{n}
\end{array}\right]\left[\begin{array}{l}
\left.\frac{\partial p}{\partial w}\right] \\
\\
\end{array}=u_{n} I+\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
u_{3} \\
H
\end{array}\right]\left[-u_{n}, n_{1}, n_{2}, n_{3}, 0\right]+(\gamma-1)\left[\begin{array}{c}
0 \\
n_{1} \\
n_{2} \\
n_{3} \\
u_{n}
\end{array}\right]\left[\frac{u^{2}}{2},-u_{1},-u_{2},-u_{3}, 1\right]\right.
\end{aligned}
$$

Note that every entry in the Jacobian matrix can be expressed in terms of the velocity components $u_{i}$ and the speed of sound $c$ since

$$
H=\frac{c^{2}}{\gamma-1}+\frac{u^{2}}{2}
$$

It may also be directly verified that

$$
f=A \boldsymbol{w}
$$

because $u_{n}$ is homogeneous of degree 0 with the consequence that $\frac{\partial u_{n}}{\partial \boldsymbol{w}}$ is orthogonal to $\boldsymbol{w}$.

$$
\frac{\partial u_{n}}{\partial \boldsymbol{w}} \boldsymbol{w}=0
$$

while $p$ is homogenous of degree 1 so that

$$
\frac{\partial p}{\partial \boldsymbol{w}} \boldsymbol{w}=p
$$

Thus

$$
A \boldsymbol{w}=u_{n} \boldsymbol{w}+\left[\begin{array}{c}
1 \\
n_{1} \\
n_{2} \\
n_{3} \\
u_{n}
\end{array}\right]\left[\frac{\partial p}{\partial \boldsymbol{w}}\right] \boldsymbol{w}=\boldsymbol{f}
$$

Also

$$
\frac{\partial \boldsymbol{f}}{\partial \boldsymbol{w}}=A+\frac{\partial A}{\partial \boldsymbol{w}} \boldsymbol{w}=A
$$

because $\frac{\partial A}{\partial \boldsymbol{w}}$ is homogeneous of degree 0 , with the consequence that $\boldsymbol{w}$ is in the null space of $\frac{\partial A}{\partial \boldsymbol{w}}$.
The special structure of the Jacobian matrix enables the direct identification of its eigenvalues and eigenvectors. Any vector in the 3 dimensional subspace orthogonal to the vectors $\frac{\partial u_{n}}{\partial \boldsymbol{w}}$ and $\frac{\partial p}{\partial \boldsymbol{w}}$ is an eigenvector corresponding to the eigenvalue $u_{n}$, which is thus a triple eigenvalue. It is easy to verify that the vectors

$$
r_{0}=\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
u_{3} \\
\frac{u^{2}}{2}
\end{array}\right] \quad r_{1}=\left[\begin{array}{c}
0 \\
0 \\
n_{3} \\
-u_{2} \\
u_{2} n_{3}-u_{3} n_{2}
\end{array}\right] \quad r_{2}=\left[\begin{array}{c}
0 \\
-n_{3} \\
0 \\
n_{1} \\
u_{3} n_{1}-u_{1} n_{3}
\end{array}\right] \quad r_{3}=\left[\begin{array}{c}
0 \\
n_{2} \\
-n_{1} \\
0 \\
u_{1} n_{2}-u_{2} n_{1}
\end{array}\right]
$$

are orthogonal to both $\frac{\partial u_{n}}{\partial \boldsymbol{w}}$ and $\frac{\partial p}{\partial \boldsymbol{w}}$. However, $r_{1}, r_{2}$ and $r_{3}$ are not independent since

$$
\sum_{k=1}^{3} n_{k} r_{k}=0
$$

Three independent eigenvectors can be obtained as

$$
v_{1}=n_{1} r_{0}+c r_{1}, v_{2}=n_{2} r_{0}+c r_{2}, v_{3}=n_{3} r_{0}+c r_{3}
$$

where $c$ is the speed of sound.
In order to verify this note that the middle three elements of $r_{k}, k=1,2,3$ are equal to $\vec{i}_{k} \times \vec{n}$, $\underset{\rightarrow}{\text { where }} \vec{i}_{k}$ is the unit vector in the $k^{t h}$ coordinate direction. Also the last element of $v_{k}$ is equal to $\vec{i}_{k} \cdot\left(\vec{n}_{k} \times \vec{u}\right)$. If the vectors $v_{k}$ are not independent they must satisfy the relation of the form

$$
\vec{v}=\sum_{k=1}^{3} \alpha_{k} \vec{v}_{k}=0
$$

for some non zero vector $\vec{\alpha}$. The first element of $\vec{v}$ is

$$
\sum_{k=1}^{3} \alpha_{k} n_{k}=\vec{\alpha} \cdot \vec{n}
$$

For the next three elements of $\vec{v}$ to be zero

$$
\sum_{k=1}^{3} \alpha_{k}\left(\vec{i}_{k} \times \vec{n}\right)=\vec{\alpha} \times \vec{n}=0
$$

which is only possible if $\vec{\alpha}$ is parallel to $\vec{n}$, so that $\vec{\alpha} \cdot \vec{n} \neq 0$.
In order to identify the remaining eigenvectors denote the column vectors in $A$ as

$$
a_{1}=\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
u_{3} \\
H
\end{array}\right] \quad a_{2}=\left[\begin{array}{c}
0 \\
n_{1} \\
n_{2} \\
n_{3} \\
u_{n}
\end{array}\right]
$$

Then

$$
\begin{aligned}
\rho \frac{\partial u_{n}}{\partial \boldsymbol{w}} a_{1} & =0, p \frac{\partial u_{n}}{\partial \boldsymbol{w}} a_{2}
\end{aligned}=1 .
$$

Now consider a vector of the form

$$
r=a_{1}+\alpha a_{2}
$$

Then

$$
\begin{align*}
A r & =u_{n}\left(a_{1}+\alpha a_{2}\right)+\alpha a_{1}+c^{2} a_{2}  \tag{1.15}\\
& =\lambda\left(a_{1}+\alpha a_{2}\right)=\lambda r
\end{align*}
$$

if

$$
u_{n}+\alpha=\lambda, \alpha u_{n}+c^{2}=\lambda
$$

which is the case if

$$
\alpha^{2}=c^{2}
$$

Thus the vectors

$$
v_{4}=a_{1}+c a_{2}, v_{5}=a_{1}-c a_{2}
$$

are the eigenvectors corresponding to the eigenvalues $u_{n}+c$ and $u_{n}-c$. Written in full

$$
v_{4}=\left[\begin{array}{c}
1 \\
u_{1}+n_{1} c \\
u_{2}+n_{2} c \\
u_{3}+n_{3} c \\
H+u_{n} c
\end{array}\right] \quad v_{5}=\left[\begin{array}{c}
1 \\
u_{1}-n_{1} c \\
u_{2}-n_{2} c \\
u_{3}-n_{3} c \\
H-u_{n} c
\end{array}\right]
$$

### 1.3 Two dimensional flow

The equations of two dimensional flow have the simpler form

$$
\frac{\partial \boldsymbol{w}}{\partial t}+\frac{\partial \boldsymbol{f}_{1}(\boldsymbol{w})}{\partial x_{1}}+\frac{\partial \boldsymbol{f}_{2}(\boldsymbol{w})}{\partial x_{2}}=0
$$

where the state and flux vectors are

$$
\boldsymbol{w}=\left[\begin{array}{c}
p \\
m_{1} \\
m_{2} \\
e
\end{array}\right], \boldsymbol{f}_{1}=u_{1} \boldsymbol{w}+p\left[\begin{array}{c}
0 \\
1 \\
0 \\
u_{1}
\end{array}\right], \boldsymbol{f}_{2}=u_{2} \boldsymbol{w}+p\left[\begin{array}{c}
0 \\
0 \\
1 \\
u_{2}
\end{array}\right]
$$

Now the flux vector normal to an edge with normal components $n_{1}$ and $n_{2}$ is

$$
\boldsymbol{f}=n_{1} \boldsymbol{f}_{1}+n_{2} \boldsymbol{f}_{2}
$$

The corresponding Jacobian matrix is

$$
\begin{aligned}
A & =\frac{\partial \boldsymbol{f}}{\partial \boldsymbol{w}}=n_{1} \frac{\partial \boldsymbol{f}_{1}}{\partial \boldsymbol{w}}+n_{2} \frac{\partial \boldsymbol{f}_{2}}{\partial \boldsymbol{w}} \\
& =u_{n} I+\rho\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
H
\end{array}\right] \frac{\partial u_{n}}{\partial \boldsymbol{w}}+\left[\begin{array}{c}
0 \\
n_{1} \\
n_{2} \\
u_{n}
\end{array}\right] \frac{\partial p}{\partial \boldsymbol{w}}
\end{aligned}
$$

The eigenvalues of $A$ are

$$
u_{n}, \quad u_{n}, \quad u_{n}+c, u_{n}-c
$$

with corresponding eigenvectors

$$
v_{1}=\left[\begin{array}{c}
1 \\
u_{1} \\
u_{2} \\
\frac{u^{2}}{2}
\end{array}\right] \quad v_{2}=\left[\begin{array}{c}
0 \\
-c n_{2} \\
c n_{1} \\
c\left(u_{2} n_{1}-u_{1} n_{2}\right)
\end{array}\right] \quad v_{3}=\left[\begin{array}{c}
1 \\
u_{1}+n_{1} c \\
u_{2}+n_{2} c \\
H+u_{n} c
\end{array}\right] \quad v_{4}=\left[\begin{array}{c}
1 \\
u_{1}-n_{1} c \\
u_{2}-n_{2} c \\
H-u_{n} c
\end{array}\right]
$$

### 1.4 One dimensional flow

The equations of one dimensional flow are further simplified to

$$
\frac{\partial \boldsymbol{w}}{\partial t}+\frac{\partial}{\partial x} \boldsymbol{f}(\boldsymbol{w})=0
$$

where the state vector is

$$
\boldsymbol{w}=\left[\begin{array}{c}
\rho \\
\rho u \\
\rho E
\end{array}\right]=\left[\begin{array}{c}
\rho \\
m \\
e
\end{array}\right]
$$

and the flux vector is

$$
\boldsymbol{f}=\left[\begin{array}{c}
\rho u \\
\rho u^{2}+p \\
\rho u H
\end{array}\right]=u \boldsymbol{w}+\left[\begin{array}{c}
0 \\
u \\
u p
\end{array}\right]
$$

Now the Jacobian matrix is

$$
A=\frac{\partial f}{\partial \boldsymbol{w}}=u I+\rho\left[\begin{array}{c}
1 \\
u \\
H
\end{array}\right] \frac{\partial u}{\partial \boldsymbol{w}}+\left[\begin{array}{l}
0 \\
1 \\
u
\end{array}\right] \frac{\partial p}{\partial \boldsymbol{w}}=u I+\left[\begin{array}{c}
1 \\
u \\
H
\end{array}\right][-u, 1,0]+(\gamma-1)\left[\begin{array}{l}
0 \\
1 \\
u
\end{array}\right]\left[\frac{u^{2}}{2},-u, 1\right]
$$

The eigenvalues of A are

$$
u, u+c, u-c
$$

with the corresponding eigenvectors

$$
v_{1}=\left[\begin{array}{c}
1 \\
u \\
\frac{u^{2}}{2}
\end{array}\right] \quad v_{2}=\left[\begin{array}{c}
1 \\
u+c \\
H+u c
\end{array}\right] \quad v_{3}=\left[\begin{array}{c}
1 \\
u-c \\
H-u c
\end{array}\right]
$$

### 1.5 Transformation to alternative sets of variables: primitive form

The quasi-linear equations of gas dynamics can be simplified in various ways by transformations to alternative sets of variables.

Consider the equations of three dimensional flow

$$
\frac{\partial \boldsymbol{w}}{\partial t}+A_{i} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0
$$

where

$$
A_{i}=\frac{\partial f}{\partial \boldsymbol{w}_{i}}
$$

Under a transformation to new set of variables $\tilde{\boldsymbol{w}}$

$$
\tilde{M} \frac{\partial \tilde{\boldsymbol{w}}}{\partial t}+A_{i} \tilde{M} \frac{\partial \tilde{\boldsymbol{w}}}{\partial x_{i}}=0
$$

where

$$
\tilde{M}=\frac{\partial \boldsymbol{w}}{\partial \tilde{\boldsymbol{w}}}
$$

Multiplying by $\frac{\partial \tilde{\boldsymbol{w}}}{\partial \boldsymbol{w}}=M^{-1}$,

$$
\frac{\partial \tilde{\boldsymbol{w}}}{\partial t}+\tilde{A}_{i} \frac{\partial \tilde{\boldsymbol{w}}}{\partial x_{i}}=0
$$

where

$$
\tilde{A}_{i}=\tilde{M}^{-1} A_{i} \tilde{M}, A_{i}=\tilde{M} \tilde{A}_{i} \tilde{M}^{-1}
$$

In the case of the primitive variables,

$$
\tilde{\boldsymbol{w}}=\left[\begin{array}{c}
\rho \\
u_{1} \\
u_{2} \\
u_{3} \\
p
\end{array}\right]
$$

we find that

$$
\begin{gathered}
\tilde{M}=\left[\begin{array}{ccccc}
1 & 0 & 0 & 0 & 0 \\
u_{1} & \rho & 0 & 0 & 0 \\
u_{2} & 0 & \rho & 0 & 0 \\
u_{3} & 0 & 0 & \rho & 0 \\
\frac{u^{2}}{2} & u_{1} & u_{2} & u_{3} & \frac{1}{\gamma-1}
\end{array}\right] \\
\tilde{M}^{-1}=\left[\begin{array}{ccccc}
1 & 0 & 0 & 0 & 0 \\
-\frac{u_{1}}{\rho} & \frac{1}{\rho} & 0 & 0 & 0 \\
-\frac{u_{2}}{\rho} & 0 & \frac{1}{\rho} & 0 & 0 \\
-\frac{u_{3}}{\rho} & 0 & 0 & \frac{1}{\rho} & 0 \\
(\gamma-1) \frac{u^{2}}{2} & -(\gamma-1) u_{1} & -(\gamma-1) u_{2} & -(\gamma-1) u_{2} & \gamma-1
\end{array}\right]
\end{gathered}
$$

and

$$
A_{1}=\left[\begin{array}{ccccc}
u_{1} & p & 0 & 0 & 0 \\
0 & u_{1} & 0 & 0 & \frac{1}{\rho} \\
0 & 0 & u_{1} & 0 & 0 \\
0 & 0 & 0 & u_{1} & 0 \\
0 & \rho c^{2} & 0 & 0 & u_{1}
\end{array}\right]
$$

$$
\begin{aligned}
& A_{2}=\left[\begin{array}{ccccc}
u_{2} & 0 & p & 0 & 0 \\
0 & u_{2} & 0 & 0 & 0 \\
0 & 0 & u_{2} & 0 & \frac{1}{\rho} \\
0 & 0 & 0 & u_{2} & 0 \\
0 & 0 & \rho c^{2} & 0 & u_{2}
\end{array}\right] \\
& A_{3}
\end{aligned}=\left[\begin{array}{ccccc}
u_{3} & 0 & 0 & p & 0 \\
0 & u_{3} & 0 & 0 & 0 \\
0 & 0 & u_{3} & 0 & 0 \\
0 & 0 & 0 & u_{3} & \frac{1}{\rho} \\
0 & 0 & 0 & \rho c^{2} & u_{3}
\end{array}\right],
$$

### 1.6 Symmetric form

Consider the equations of one dimensional flow in primitive variables

$$
\begin{gathered}
\frac{\partial \rho}{\partial t}+u \frac{\partial \rho}{\partial x}+\rho \frac{\partial u}{\partial x}=0 \\
\frac{\partial u}{\partial t}+u \frac{\partial u}{\partial x}+\frac{1}{\rho} \frac{\partial p}{\partial x}=0 \\
\frac{\partial p}{\partial t}+\rho c^{2} \frac{\partial u}{\partial x}+u \frac{\partial p}{\partial x}=0
\end{gathered}
$$

Then subtracting the first equation multiplied by $c^{2}$ from the third equation, we find that

$$
\frac{\partial p}{\partial t}-c^{2} \frac{\partial \rho}{\partial t}+u\left(\frac{\partial p}{\partial x}-c^{2} \frac{\partial \rho}{\partial x}\right)=0
$$

This is equivalent to a statement that the entropy

$$
S=\log \left(\frac{p}{\rho^{\gamma}}\right)=\log p-\gamma \log \rho
$$

is constant since

$$
d S=\frac{d p}{p}-\gamma \frac{d \rho}{\rho}=\frac{1}{p}\left(d p-c^{2} d \rho\right)
$$

If the entropy is constant, then

$$
d p=c^{2} d \rho
$$

With this substitution the first equation becomes,

$$
\frac{1}{c^{2}} \frac{\partial p}{\partial t}+\frac{u}{c^{2}} \frac{\partial p}{\partial x}+\rho \frac{\partial u}{\partial x}=0
$$

and now the first two equations can be rescaled as

$$
\begin{gathered}
\frac{1}{\rho c} \frac{\partial p}{\partial t}+\frac{u}{\rho c} \frac{\partial p}{\partial x}+c \frac{\partial u}{\partial x}=0 \\
\frac{\partial u}{\partial t}+\frac{c}{\rho c} \frac{\partial p}{\partial x}+u \frac{\partial u}{\partial x}=0
\end{gathered}
$$

Thus if we write the equations in terms of the differential variables

$$
d \overline{\boldsymbol{w}}=\left[\begin{array}{c}
\frac{d p}{\rho c} \\
d u \\
d p-c^{2} d \rho
\end{array}\right]
$$

we obtain the symmetric form

$$
\frac{\partial \overline{\boldsymbol{w}}}{\partial t}+\bar{A} \frac{\partial \overline{\boldsymbol{w}}}{\partial x}=0
$$

where

$$
\bar{A}=\left[\begin{array}{lll}
u & c & 0 \\
c & u & 0 \\
0 & 0 & u
\end{array}\right]
$$

These transformations can be generalized to the equations of three dimensional flow. A convenient scaling is to set

$$
d \overline{\boldsymbol{w}}=\left[\begin{array}{c}
\frac{d p}{c^{2}}  \tag{1.16}\\
\frac{\rho}{c} d u_{1} \\
\frac{\rho}{c} d u_{2} \\
\frac{\rho}{c} d u_{3} \\
\frac{d p}{c^{2}}-d \rho
\end{array}\right]
$$

This eliminates the density from the transformation matrices $\bar{M}=\frac{d \boldsymbol{w}}{d \overline{\boldsymbol{w}}}$ and $\bar{M}^{-1}=\frac{d \overline{\boldsymbol{w}}}{d \boldsymbol{w}}$. The equations now take the form

$$
\begin{equation*}
\frac{\partial \overline{\boldsymbol{w}}}{\partial t}+\bar{A}_{i} \frac{\partial \overline{\boldsymbol{w}}}{\partial x_{i}}=0 \tag{1.17}
\end{equation*}
$$

where the transformed Jacobian matrices

$$
\begin{equation*}
\bar{A}_{i}=\bar{M}^{-1} A_{i} \bar{M} \tag{1.18}
\end{equation*}
$$

are simultaneously symmetrized as

$$
\begin{align*}
& \bar{A}_{1}=\left[\begin{array}{ccccc}
u_{1} & c & 0 & 0 & 0 \\
c & u_{1} & 0 & 0 & 0 \\
0 & 0 & u_{1} & 0 & 0 \\
0 & 0 & 0 & u_{1} & 0 \\
0 & 0 & 0 & 0 & u_{1}
\end{array}\right]  \tag{1.19}\\
& \bar{A}_{2}=\left[\begin{array}{ccccc}
u_{2} & 0 & c & 0 & 0 \\
0 & u_{2} & 0 & 0 & 0 \\
c & 0 & u_{2} & 0 & 0 \\
0 & 0 & 0 & u_{2} & 0 \\
0 & 0 & 0 & 0 & u_{2}
\end{array}\right]  \tag{1.20}\\
& \bar{A}_{3}=\left[\begin{array}{ccccc}
u_{3} & 0 & 0 & c & 0 \\
0 & u_{3} & 0 & 0 & 0 \\
0 & 0 & u_{3} & 0 & 0 \\
c & 0 & 0 & u_{3} & 0 \\
0 & 0 & 0 & 0 & u_{3}
\end{array}\right] \tag{1.21}
\end{align*}
$$

while

$$
\bar{M}=\frac{\partial \boldsymbol{w}}{\partial \overline{\boldsymbol{w}}}=\left[\begin{array}{ccccc}
1 & 0 & 0 & 0 & -1  \tag{1.22}\\
u_{1} & c & 0 & 0 & -u_{1} \\
u_{2} & 0 & c & 0 & -u_{2} \\
u_{3} & 0 & 0 & c & -u_{3} \\
H & c u_{1} & c u_{2} & c u_{3} & -\frac{u^{2}}{2}
\end{array}\right]
$$

and

$$
\bar{M}^{-1}=\frac{\partial \overline{\boldsymbol{w}}}{\partial \boldsymbol{w}}=\left[\begin{array}{ccccc}
\bar{\gamma} \frac{u^{2}}{2} & -\bar{\gamma} u_{1} & -\bar{\gamma} u_{2} & -\bar{\gamma} u_{3} & -\bar{\gamma}  \tag{1.23}\\
-\frac{u_{1}}{c} & \frac{1}{c} & 0 & 0 & 0 \\
-\frac{u_{c}}{c} & 0 & \frac{1}{c} & 0 & 0 \\
-\frac{u_{3}}{c} & 0 & 0 & \frac{1}{c} & 0 \\
\bar{\gamma}\left(u^{2}-H\right) & -\bar{\gamma} u_{1} & -\bar{\gamma} u_{2} & -\bar{\gamma} u_{3} & \bar{\gamma}
\end{array}\right]
$$

where

$$
\begin{equation*}
\bar{\gamma}=\frac{\gamma-1}{c^{2}} \tag{1.24}
\end{equation*}
$$

The combined Jacobian matrix

$$
\begin{equation*}
A=n_{i} A_{i} \tag{1.25}
\end{equation*}
$$

can now be decomposed as

$$
\begin{equation*}
A=n_{i} \bar{M} \bar{A}_{i} \bar{M}^{-1} \tag{1.26}
\end{equation*}
$$

Corresponding to the fact that $\bar{A}$ is symmetric one can find a set of orthogonal eigenvectors, which may be normalized to unit length. Then one can express

$$
\bar{A}=\bar{V} \Lambda \bar{V}^{-1}
$$

where the diagonal matrix $\Lambda$ contains the eigenvalues $u_{n}, u, u_{n}, u_{n}+c$ and $u_{n}-c$ as its elements. The matrix $\bar{V}$ containing the corresponding eigenvectors as its columns is

$$
\bar{V}=\left[\begin{array}{ccccc}
\frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 & 0 & 0  \tag{1.27}\\
\frac{n 1}{\sqrt{2}} & \frac{n 1}{\sqrt{2}} & 0 & -n_{3} & n_{2} \\
\frac{n 2}{\sqrt{2}} & \frac{n_{2}}{\sqrt{2}} & n_{3} & 0 & -n_{1} \\
\frac{n 3}{\sqrt{2}} & \frac{n 3}{\sqrt{2}} & -n_{2} & n_{1} & 0 \\
0 & 0 & 1 & n_{2} & n_{3}
\end{array}\right]
$$

and $\bar{V}^{-1}=\bar{V}^{T}$. The Jacobian matrix can now be expressed as

$$
\begin{equation*}
A=M \Lambda M^{-1} \tag{1.28}
\end{equation*}
$$

where

$$
\begin{equation*}
M=\bar{M} \bar{V}, M^{-1}=\bar{V}^{T} \bar{M}^{-1} \tag{1.29}
\end{equation*}
$$

This decomposition is often useful.
Since

$$
c^{2}=\frac{d p}{d \rho}
$$

it follows that in isentropic flow

$$
2 c d c=\frac{\gamma}{\rho} d \rho-\frac{\gamma p}{\rho^{2}} d \rho=\frac{\gamma}{\rho} d \rho-c^{2} d \rho=\frac{\gamma-1}{\rho} d p
$$

Thus

$$
\frac{d p}{\rho c}=\frac{2}{\gamma-1} d c
$$

so the equations can also be expressed in the same form, equations $(1.17),(1.18),(1.19),(1.20)$ and (1.21) for the variables

$$
\overline{\boldsymbol{w}}=\left[\begin{array}{c}
\frac{2 c}{\gamma-1} \\
u_{1} \\
u_{2} \\
u_{3} \\
S
\end{array}\right]
$$

with the transformation matrices

$$
\bar{M}=\frac{d \boldsymbol{w}}{d \overline{\boldsymbol{w}}}=\left[\begin{array}{ccccc}
\frac{\rho}{c} & 0 & 0 & 0 & -\frac{p}{c^{2}} \\
\frac{\rho u_{1}}{c} & \rho & 0 & 0 & -\frac{p u_{1}}{c^{2}} \\
\frac{\rho u_{2}}{c} & 0 & \rho & 0 & -\frac{p u_{2}}{c^{2}} \\
\frac{\rho u_{3}}{c} & 0 & 0 & \rho & -\frac{p u_{3}}{c^{2}} \\
\frac{\rho H}{c} & \rho u_{1} & \rho u_{2} & \rho u_{3} & -\rho \frac{p}{c} \frac{u^{2}}{2}
\end{array}\right]
$$

and

$$
\bar{M}^{-1}=\frac{d \overline{\boldsymbol{w}}}{d \boldsymbol{w}}=\left[\begin{array}{ccccc}
\bar{\gamma} \frac{u^{2}}{2} & -\bar{\gamma} u_{1} & -\bar{\gamma} u_{2} & -\bar{\gamma} u_{3} & \bar{\gamma} \\
-\frac{u_{1}}{\rho} & \frac{1}{\rho} & 0 & 0 & 0 \\
-\frac{u_{2}}{\rho} & 0 & \frac{1}{\rho} & 0 & 0 \\
-\frac{u_{3}}{\rho} & 0 & 0 & \frac{1}{\rho} & 0 \\
\overline{\frac{\gamma}{p}}\left(u^{2}-H\right) & -\frac{\bar{\gamma} u_{1}}{p} & -\frac{\bar{\gamma} u_{2}}{p} & -\frac{\bar{\gamma} u_{3}}{p} & \frac{\bar{\gamma}}{p}
\end{array}\right]
$$

where

$$
\bar{\gamma}=\frac{\gamma-1}{\rho c}
$$

### 1.7 Riemann invariants

In the case of one dimensional isentropic flow these equations reduce to

$$
\begin{gathered}
\frac{2}{\gamma-1} \frac{\partial c}{\partial t}+\frac{2 u}{\gamma-1} \frac{\partial c}{\partial x}+c \frac{\partial u}{\partial x}=0 \\
\frac{\partial u}{\partial t}+\frac{2 c}{\gamma-1} \frac{\partial c}{\partial x}+u \frac{\partial u}{\partial x}=0 \\
\frac{\partial S}{\partial t}+u \frac{\partial S}{\partial x}=0
\end{gathered}
$$

Now the first two equations can be added and subtracted to yield

$$
\frac{\partial R^{+}}{\partial t}+(u+c) \frac{\partial R^{+}}{\partial x}=0
$$

and

$$
\frac{\partial R^{-}}{\partial t}+(u-c) \frac{\partial R^{-}}{\partial x}=0
$$

where $R^{+}$and $R^{-}$are the Riemann invariants

$$
R^{+}=u+\frac{2 c}{\gamma-1}, R^{-}=u-\frac{2 c}{\gamma-1}
$$

which remain constant as they are transported at the wave speeds $u+c$ and $u-c$. The Riemann invariants prove to be useful in the formulation of far field boundary conditions designed to minimize wave reflection.

### 1.8 Symmetric hyperbolic form

Equation (1.17) can be written in terms of the conservative variables as

$$
\frac{\partial \overline{\boldsymbol{w}}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial t}+\bar{A}_{i} \frac{\partial \overline{\boldsymbol{w}}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0
$$

Here $\frac{\partial \bar{w}}{\partial \boldsymbol{w}}=\bar{M}^{-1}$. Now multiplying by $\bar{M}^{T^{-1}}$ the equation is reduced to the symmetric hyperbolic form

$$
\begin{equation*}
Q \frac{\partial \boldsymbol{w}}{\partial t}+\hat{A}_{i} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0 \tag{1.30}
\end{equation*}
$$

where $Q$ is symmetric and positive definite and the matrices $\hat{A}_{i}$ are symmetric

$$
\begin{equation*}
Q=\left(M M^{T}\right)^{-1}, \hat{A}_{i}=\bar{M}^{T^{-1}} \bar{A}_{i} \bar{M}^{-1} \tag{1.31}
\end{equation*}
$$

This form could alternatively be derived by multiplying the conservative form (1.2) of the equations by $Q$. Then

$$
\begin{aligned}
Q \frac{\partial \boldsymbol{f}_{i}}{\partial x_{i}} & =\bar{M}^{T^{-1}} \bar{M}^{-1} \frac{\partial \boldsymbol{f}}{\partial x_{i}} \\
& =\bar{M}^{T^{-1}} \bar{M}^{-1} A_{i} \frac{\partial \boldsymbol{w}}{\partial x_{i}} \\
& =\bar{M}^{T^{-1}} \bar{M}^{-1} M \bar{A}_{i} \bar{M}^{-1} \frac{\partial \boldsymbol{w}}{\partial x_{i}} \\
& =\bar{M}^{T^{-1}} A_{i} \bar{M}^{-1} \frac{\partial \boldsymbol{w}}{\partial x_{i}}
\end{aligned}
$$

A symmetric hyperbolic form can actually be obtained for any choice of the dependent variables. For example equation (1.17) could be written in terms of the primitive variables as

$$
N \frac{\partial \tilde{\boldsymbol{w}}}{\partial t}+\bar{A}_{i} N \frac{\partial \tilde{\boldsymbol{w}}}{\partial x_{i}}=0
$$

where

$$
N=\frac{\partial \overline{\boldsymbol{w}}}{\partial \tilde{\boldsymbol{w}}}=\left[\begin{array}{ccccc}
0 & 0 & 0 & 0 & \frac{1}{c^{2}} \\
0 & \frac{\rho}{c} & 0 & 0 & 0 \\
0 & 0 & \frac{\rho}{c} & 0 & 0 \\
0 & 0 & 0 & \frac{\rho}{c} & 0 \\
-1 & 0 & 0 & 0 & \frac{1}{c^{2}}
\end{array}\right]
$$

Then multiplying by $N^{T}$ we obtain the symmetric hyperbolic form

$$
N^{T} N \frac{\partial \tilde{\boldsymbol{w}}}{\partial t}+N^{T} \bar{A}_{i} N \frac{\partial \tilde{\boldsymbol{w}}}{\partial x_{i}}=0
$$

### 1.9 Entropy variables

A particular choice of variables which symmetrizes the equations can be derived from functions of the entropy

$$
\begin{equation*}
S=\log \left(\frac{p}{\rho \gamma}\right)=\log p-\gamma \log \rho \tag{1.32}
\end{equation*}
$$

The last equation of (1.17) is equivalent to the statement that

$$
\rho \frac{\partial S}{\partial t}+\rho u_{i} \frac{\partial S}{\partial x_{i}}=0
$$

which can be combined with the mass conservation equation multiplied by $S$

$$
S \frac{\partial \rho}{\partial t}+S \frac{\partial}{\partial x_{i}}\left(\rho u_{i}\right)=0
$$

to yield the entropy conservation law

$$
\begin{equation*}
\frac{\partial(\rho S)}{\partial t}+\frac{\partial\left(\rho u_{i} S\right)}{\partial x_{i}}=0 \tag{1.33}
\end{equation*}
$$

This is a special case of a generalized entropy function defined as follows. Given a system of conservation laws

$$
\begin{equation*}
\frac{\partial \boldsymbol{w}}{\partial t}+\frac{\partial}{\partial x_{i}} \boldsymbol{f}_{i}(\boldsymbol{w})=0 \tag{1.34}
\end{equation*}
$$

Suppose that we can find a scalar function $U(\boldsymbol{w})$ such that

$$
\begin{equation*}
\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}}=\frac{\partial G_{i}}{\partial \boldsymbol{w}} \tag{1.35}
\end{equation*}
$$

and $U(\boldsymbol{w})$ is a convex function of $\boldsymbol{w}$. Then $U(\boldsymbol{w})$ is an entropy function with an entropy flux $G_{i}(\boldsymbol{w})$ since multiplying equation (1.34) by $\frac{\partial U}{\partial w}$ we obtain

$$
\begin{equation*}
\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial t}+\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0 \tag{1.36}
\end{equation*}
$$

and using (1.35) this equivalent to

$$
\frac{\partial U(\boldsymbol{w})}{\partial t}+\frac{\partial G_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0
$$

which is in turn equivalent to the generalized entropy conservation law

$$
\begin{equation*}
\frac{\partial U(\boldsymbol{w})}{\partial t}+\frac{\partial}{\partial x_{i}} G_{i}=0 \tag{1.37}
\end{equation*}
$$

Now introduce dependent variables

$$
\begin{equation*}
v^{T}=\frac{\partial U}{\partial \boldsymbol{w}} \tag{1.38}
\end{equation*}
$$

Then the equations are symmetrized. Define the scalar functions

$$
\begin{equation*}
Q(v)=v^{T} \boldsymbol{w}-U(\boldsymbol{w}) \tag{1.39}
\end{equation*}
$$

and

$$
\begin{equation*}
R_{i}(v)=v^{T} \boldsymbol{f}_{i}-G_{i}(\boldsymbol{w}) \tag{1.40}
\end{equation*}
$$

Then

$$
\frac{\partial Q}{\partial v}=\boldsymbol{w}^{T}+v^{T} \frac{\partial \boldsymbol{w}}{\partial v}-\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial v}=\boldsymbol{w}^{T}
$$

and

$$
\frac{\partial R_{i}}{\partial v}=\boldsymbol{f}_{i}^{T}+v^{T} \frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial v}-\frac{\partial G_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial v}=\boldsymbol{f}_{i}^{T}
$$

Hence $\frac{\partial w}{\partial v}$ is symmetric.

$$
\frac{\partial \boldsymbol{w}_{j}}{\partial v_{k}}=\frac{\partial^{2} Q}{\partial v_{j} \partial v_{k}}=\frac{\partial \boldsymbol{w}_{k}}{\partial v_{j}}
$$

and $\frac{\partial f_{i}}{\partial v}$ is symmetric

$$
\frac{\partial \boldsymbol{f}_{i j}}{\partial v_{k}}=\frac{\partial^{2} R_{i}}{\partial v_{j} \partial v_{k}}=\frac{\partial \boldsymbol{f}_{i k}}{\partial v_{i}}
$$

and correspondingly

$$
\frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial v}=A_{i} \frac{\partial \boldsymbol{w}}{\partial v}
$$

is symmetric. Thus the conservation law (1.34) is converted to the symmetric form

$$
\begin{equation*}
\frac{\partial \boldsymbol{w}}{\partial v} \frac{\partial v}{\partial t}+A_{i} \frac{\partial \boldsymbol{w}}{\partial v} \frac{\partial v}{\partial x_{i}}=0 \tag{1.41}
\end{equation*}
$$

Multiplying the conservation law (1.34) by $\frac{\partial w}{\partial v}$ from the left produces the alternative symmetric form

$$
\begin{equation*}
\frac{\partial v}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial t}+\frac{\partial v}{\partial \boldsymbol{w}} A_{i} \frac{\partial \boldsymbol{w}}{\partial x_{i}}=0 \tag{1.42}
\end{equation*}
$$

since

$$
\frac{\partial v}{\partial \boldsymbol{w}} A_{i}=\frac{\partial v}{\partial \boldsymbol{w}}\left(A_{i} \frac{\partial \boldsymbol{w}}{\partial v}\right) \frac{\partial v}{\partial \boldsymbol{w}}
$$

Moreover, multiplying (1.34) by $v^{T}$ recovers equation (1.36) and the corresponding entropy conservation law (1.37).

Conversely, if $U(\boldsymbol{w})$ satisfies the conservation law (1.37), then

$$
\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial t}=-\frac{\partial G_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial x_{i}}
$$

but

$$
\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial t}=-\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial x_{i}}
$$

These can both hold for all $\frac{\partial w}{\partial x_{i}}$ only if

$$
\frac{\partial U}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{f}_{i}}{\partial \boldsymbol{w}}=\frac{\partial G_{i}}{\partial \boldsymbol{w}}
$$

so $u(\boldsymbol{w})$ is an entropy function if it is a convex function of $\boldsymbol{w}$.

The property of convexity enables $U(\boldsymbol{w})$ to be treated as a generalized energy, since $u$ becomes unbounded if $\boldsymbol{w}$ becomes unbounded. It can be shown that $-\rho S$ is a convex function of $\boldsymbol{w}$. Accordingly it follows from the entropy equation (1.33) that we can take

$$
\begin{equation*}
U(\boldsymbol{w})=\frac{\rho S}{\gamma-1}=\frac{\gamma}{\gamma-1} \rho \log \rho-\frac{\rho}{\gamma-1} \log p \tag{1.43}
\end{equation*}
$$

as a generalized entropy function where the scaling factor $\frac{1}{\gamma-1}$ is introduced to simplify the resulting expressions.

Using the formula

$$
\frac{\partial p}{\partial \boldsymbol{w}}=(\gamma-1)\left[\frac{u^{2}}{2},-u_{1},-u_{2},-u_{3}, 1\right]
$$

we find that

$$
\begin{equation*}
v^{T}=\frac{\partial u}{\partial \boldsymbol{w}}=\left[\frac{\gamma-S}{\gamma-1} \frac{\rho}{p} \frac{u^{2}}{2}, \frac{\rho u_{1}}{p}, \frac{\rho u_{2}}{p}, \frac{\rho u_{3}}{p}, \frac{-\rho}{p}\right] \tag{1.44}
\end{equation*}
$$

where

$$
u^{2}=u_{i} u_{i}
$$

The reverse transformation is

$$
\begin{equation*}
\boldsymbol{w}^{T}=p\left[-v_{5}, v_{2}, v_{3}, v_{4},\left(1-\frac{1}{2}\left(v_{2}^{2}+v_{3}^{2}+v_{4}^{2}\right) v_{5}\right]\right. \tag{1.45}
\end{equation*}
$$

Now the scalar functions $Q(v)$ and $R_{i}(v)$ defined by equations (1.39) and (1.40) reduce to

$$
\begin{equation*}
Q(v)=\rho, R_{i}(v)=m_{i}=\rho u_{i} \tag{1.46}
\end{equation*}
$$

Accordingly

$$
\boldsymbol{w}^{T}=\frac{\partial p}{\partial v}=\frac{\partial p}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial v}
$$

and

$$
\boldsymbol{f}_{i}^{T}=\frac{\partial m_{i}}{\partial v}=\frac{\partial m_{i}}{\partial \boldsymbol{w}} \frac{\partial \boldsymbol{w}}{\partial v}
$$

where

$$
\frac{\partial \rho}{\partial \boldsymbol{w}}=\left[\begin{array}{lllll}
1 & 0 & 0 & 0 & 0
\end{array}\right]
$$

and

$$
\frac{\partial m_{i}}{\partial \boldsymbol{w}}=\left[\begin{array}{llll}
0 & \delta_{i 1} & \delta_{i 2} & \delta_{i 3}
\end{array}\right]
$$

Thus the first four rows of $\frac{\partial \boldsymbol{w}}{\partial v}$ consist of $\boldsymbol{w}^{T}, \boldsymbol{f}_{1}^{T}, \boldsymbol{f}_{2}^{T}$ and $\boldsymbol{f}_{3}^{T}$, and correspondingly since $\frac{\partial \boldsymbol{w}}{\partial v}$ is symmetric its first four columns are $\boldsymbol{w}, \boldsymbol{f}_{1}, \boldsymbol{f}_{2}$ and $\boldsymbol{f}_{3}$. The only remaining element is $\frac{\partial \boldsymbol{w}_{5}}{\partial v_{5}}$. A direct calculation reveals that

$$
\frac{\partial \boldsymbol{w}_{5}}{\partial v_{5}}=\rho E^{2}+\left(E+\frac{u^{2}}{2}\right) p
$$

Thus

$$
\begin{aligned}
\frac{\partial \boldsymbol{w}}{\partial v} & =\left[\begin{array}{ccccc}
\rho & \rho u_{1} & \rho u_{2} & \rho u_{3} & \rho E \\
\rho u_{1} & \rho u_{1}^{2}+p & \rho u_{2} u_{1} & \rho u_{3} u_{1} & \rho u_{1} E+u_{1} p \\
\rho u_{2} & \rho u_{1} u_{2} & \rho u_{2}^{2}+p & \rho u_{3} u_{2} & \rho u_{2} E+u_{2} p \\
\rho u_{3} & \rho u_{1} u_{3} & \rho u_{2} u_{3} & \rho u_{3}^{2}+p & \rho u_{3} E+u_{3} p \\
\rho E & \rho u_{1} E+u_{1} p & \rho u_{2} E+u_{2} p & \rho u_{3} E+u_{3} p & \rho E^{2}+\left(E+\frac{u^{2}}{2}\right) p
\end{array}\right] \\
& =\frac{1}{\rho} \boldsymbol{w} \boldsymbol{w}^{T}+p\left[\begin{array}{ccccc}
0 & 0 & 0 & 0 & 0 \\
0 & 1 & 0 & 0 & u_{1} \\
0 & 0 & 1 & 0 & u_{2} \\
0 & 0 & 0 & 1 & u_{3} \\
0 & u_{1} & u_{2} & u_{3} & E+\frac{u^{2}}{2}
\end{array}\right]
\end{aligned}
$$

